### <sup>1</sup> Hamiltonian dynamics of thermostated systems: Two-temperature <sup>2</sup> heat-conducting $\phi^4$ chains

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7 We consider and compare four Hamiltonian formulations of thermostated mechanics, three of them 8 kinetic, and the other one configurational. Though all four approaches "work" at equilibrium, their application to many-body nonequilibrium simulations can fail to provide a proper flow of heat. All 9 the Hamiltonian formulations considered here are applied to the same prototypical two-temperature 10 " $\phi^{4}$ " model of a heat-conducting chain. This model incorporates nearest-neighbor Hooke's-Law 11 interactions plus a quartic tethering potential. Physically correct results, obtained with the isokinetic 12 Gaussian and Nosé-Hoover thermostats, are compared with two other Hamiltonian results. The latter 13 results, based on constrained Hamiltonian thermostats, fail to model correctly the flow of heat. 14

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### **17 I. INTRODUCTION**

Computational "thermostats" arose as a means for con-18 19 trolling numerical simulations of both equilibrium and non-20 equilibrium stationary states. Without thermostats systems 21 driven away from equilibrium heat up. With thermostats the 22 heat generated by irreversible processes can be steadily re-23 moved, making it possible to simulate nonequilibrium steady 24 states.<sup>1,2</sup> Because most Hamiltonian-based mechanics prob-25 lems conserve energy, novel nonHamiltonian ideas are typi-26 cally required when thermostats are to be included. Never-27 theless, several approaches to Hamiltonian thermostats have 28 been developed. Here we consider various approaches pio-29 neered by Ashurst, Dettmann, Evans, Hoover, Leete, Mor-**30** riss, Nosé, and Woodcock over a span of about 25 years.<sup>1–10</sup> The simplest thermostat type maintains a (nearly) con-31 32 stant kinetic energy by "rescaling" the velocities at the end of 33 each computational time step. For a more elegant, but 34 equivalent, continuous approach to rescaling, see the "Gauss-35 ian isokinetic" method described in Sec. II B and illustrated 36 in Sec. VII. For # Cartesian degrees of freedom the kinetic-**37** theory relation

38 
$$K = \sum_{n=1}^{\infty} \frac{m\dot{q}^2}{2} = \sum_{n=1}^{\infty} \frac{p^2}{2m} = \# \frac{kT_{\rm kin}}{2},$$

**39** defines the kinetic temperature  $T_{kin}$ . Thermostats based on **40**  $T_{kin}$  can be applied to an entire system, or separately to two **41** or more subsystems. With this *ad hoc* rescaling approach

42 
$$p \rightarrow p_0 \sqrt{(K/K_0)},$$

**43** there is no difficulty in maintaining several temperatures in **44** specified parts of a single nonequilibrium system, as did **45** Ashurst in his Ph.D. thesis research.<sup>11</sup> The kinetic theory and the Gibbs-Boltzmann develop- 46 ment of thermodynamics based on ideal-gas thermometry 47 certainly suggest that the kinetic temperature,  $T_{\rm kin}$ , is both 48 the simplest and the most fundamental of the many possible 49 "temperature" types. There are alternatives. Among the many 50 possible temperatures based on particle coordinates, rather 51 than velocities, the simplest—on which we focus here—is 52 based on the mean-squared forces,  $\langle F^2 \rangle$ . This definition of a 53 configurational temperature is derived directly from the po- 54 tential energy  $\Phi$  and its space derivatives 55

$$kT = \langle (\nabla \Phi)^2 \rangle / \langle \nabla^2 \Phi \rangle = - \langle F^2 \rangle / \langle \nabla F \rangle.$$
 56

This somewhat cumbersome relation appears as an aside in 57 Landau and Lifshitz' classic 1958 text.<sup>12</sup> It was also recently 58 rediscovered by Rugh, and has subsequently been much 59 discussed.<sup>13–16</sup> This configurational representation of tem- 60 perature follows most simply from an integration by parts, in 61 the canonical ensemble: 62

$$\int dq \nabla^2 \Phi = \nabla \Phi; \ (-d/dq) e^{-\Phi/kT}$$
<sup>63</sup>

=

$$= (\nabla \Phi/kT)e^{-\Phi/kT} \to \int dq \nabla^2 \Phi e^{-\Phi/kT}$$
 64

$$= (1/kT) \int dq (\nabla \Phi)^2 e^{-\Phi/kT}.$$
65

About ten years after Woodcock and Ashurst's introduction 66 of the rescaling isokinetic thermostat in the early 1970s, 67 Hoover and Evans discovered that the isokinetic velocity- 68 rescaling equations of motion can be derived from Gauss' 69 Principle of Least Constraint.<sup>5</sup> Fifteen years after this discov- 70 ery, Dettmann and Morriss<sup>4</sup> found a straightforward Hamil- 71 tonian approach to these same motion equations. Previously, 72 Leete and Hoover had derived a Hamiltonian formulation 73 which likewise maintained the velocity-based kinetic energy 74  $K({\dot{q}})$  constant.<sup>6,8</sup> 75

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A Hamiltonian generating the Gauss' principle isokinetic requations of motion was not found until 1996.<sup>4</sup> In the meantime Nosé had discovered his canonical-ensemble thermory stated dynamics, which is related to two rather different an Hamiltonians.<sup>7,9</sup> The purpose of the present work is twofold. Hamiltonians.<sup>7,9</sup> The purpose of the Present work is twofold. First, we explore the relation of the Hoover-Leete kinetic thermostat work to its successors. A 2007 literature search shows only a single reference to it.<sup>6</sup> We also explore the consequences of a "Landau-Lifshitz" configurational thermostat based on the canonical-ensemble definition rediscovered by Rugh with a relatively complicated measure-based approach using the *microcanonical* ensemble. Recent pedagogical reviews of this configurational-temperature work can be found in Refs. 13–16.

### 90 II. THE NOSÉ-HOOVER THERMOSTAT

91 The Nosé-Hoover equations of motion<sup>1,9</sup> are a standard 92 approach to carrying out canonical-ensemble dynamics at a 93 particular temperature  $T_0$ . This approach uses *feedback*, con-94 trolling the fluctuating kinetic energy with a "friction coeffi-95 cient" or "control variable"  $\zeta$ . The frictional feedback force 96 is  $-\zeta p$ . The time rate of change of the friction coefficient is 97 based on the current value of the kinetic energy K(t) relative 98 to the desired mean value  $K_0(T_0)$ . The equations of motion 99 also include a characteristic time  $\tau$  which determines the 100 time-range over which the feedback acts

101 
$$\{\dot{q} = p/m; \quad \dot{p} = F - \zeta p\}; \quad \dot{\zeta} = [(K/K_0) - 1]/\tau^2.$$

102 It is quite interesting to see that a canonical-ensemble 103 approach<sup>1</sup> to a Nosé-Hoover-type *configurational* thermostat 104 led Braga and Travis<sup>14</sup> to the motion equations (for a har-105 monic oscillator with unit mass, force constant, and tempera-106 ture)

**107** 
$$\dot{q} = p - \zeta q; \quad \dot{p} = -q; \quad \dot{\zeta} = [q^2 - 1]/\tau^2.$$

108 The substitution  $(+q, +p, +\zeta, +t) \rightarrow (-p, -q, -\zeta, -t)$  gives

109 
$$\dot{p} = -q - \zeta p; \quad \dot{q} = +p; \quad \dot{\zeta} = [p^2 - 1]/\tau^2,$$

**110** *exactly* the same as the Nosé-Hoover equations for the *ki*-**111** *netic* thermostating of an oscillator!

112 In most cases, including our  $\phi^4$  model studies described 113 in detail in Secs. V–IX, a reasonable choice of  $\tau$  corresponds 114 to a typical collision time or vibration time. Such a choice 115 typically provides the Gibbs' equilibrium canonical phase-116 space distribution with a Gaussian distribution for the fric-117 tion coefficient  $\zeta$ 

118 
$$f(q,p,\zeta) = e^{-\Phi/kT} e^{-K/kT} e^{-\#\zeta^2 \tau^2/2}.$$

119 Evidently the amplitude of the fluctuations of  $\zeta$  varies as 120  $\sqrt{1/\#\tau^2}$ , where # is the number of degrees of freedom. Thus 121 its effect on the dynamics vanishes in the large-system limit. 122 Nonequilibrium situations are quite different. By using 123 two or more different friction coefficients (or "thermostat 124 variables") temperature differences can be established, lead-125 ing to heat flow. Then the corresponding phase-space distri-126 butions are no longer smooth and Gibbsian, but instead be-127 come fractal, with the underlying phase-space trajectories satisfying the second law of thermodynamics infinitely more **128** probable than those violating the law.<sup>2</sup> In the equilibrium **129** case there is a close connection of Nosé-Hoover mechanics **130** to Hamiltonian mechanics. We describe this connection next. **131** 

### A. The Dettmann-Nosé Hamiltonian 132

In 1984 Nosé discovered a Hamiltonian consistent with 133 Gibbs' constant-temperature canonical ensemble 134

$$f(q,p) \propto \exp(-\mathcal{H}_{Nos\acute{e}}/kT).$$
 135

He was able to derive the Nosé-Hoover equations of motion 136 described above from his Hamiltonian by an artificial "time 137 scaling." 138

In July of 1996 Dettmann<sup>7</sup> discovered a simpler ap- 139 proach to the Nosé-Hoover equations of motion. Dettmann's 140 *vanishing* Hamiltonian is 141

$$\mathcal{H}_{\text{Dettmann}} \equiv s\mathcal{H}_{\text{Nosé}} = \sum \frac{p^2}{2ms} + s \left[ \Phi + \frac{p_s^2}{2M} + \#kT \ln s \right] \equiv 0.$$
 142

Provided that Dettmann's Hamiltonian is set equal to the 143 special value, zero, Nosé's time-scaling variable *s* can be 144 eliminated. Then the Nosé-Hoover equations of motion result 145 *without the need for any time scaling*. It needs to be empha-146 sized that this  $\mathcal{H}=0$  trick does *not* work for the prototypical 147 *nonequilibrium* case of a system with two different tempera-148 tures. In a two-temperature system the two different vari-149 ables  $s_{\text{Hot/Cold}}$  are necessarily coupled and cannot both be 150 eliminated, so that Dettmann's  $\mathcal{H}=0$  trick does not work. 151

There is a special case of Nosé-Hoover dynamics that 152 deserves special mention, the "isokinetic" case in which the 153 temperature is constant. This corresponds to the choice  $\tau$  154  $\rightarrow 0$ . For isokinetic mechanics Dettmann and Morris discov-155 ered the special Hamiltonian detailed next. 156

### B. The Dettmann-Morriss isokinetic Hamiltonian 157

Soon after Dettmann discovered a Hamiltonian route to 158 the Nosé-Hoover motion equations, Dettmann and Morriss<sup>4</sup> 159 discovered a related approach to isokinetic (constant tem- 160 perature, provided that temperature is defined by the kinetic 161 energy) Hamiltonian dynamics 162

$$\mathcal{H}_{\rm DM} = K e^{+\Phi/2K_0} - K_0 e^{-\Phi/2K_0} \equiv 0; \quad K \equiv \sum \frac{p^2}{2m}.$$
 (63)

As usual, the equations of motion follow by differentiation 164

$$\{\dot{q} = \partial \mathcal{H}/\partial p = p e^{+\Phi/2K_0}; \quad \dot{p} = -\partial \mathcal{H}/\partial q = F e^{-\Phi/2K_0}\}.$$
 165

The accelerations which follow, using the identity

$$K/K_0 = e^{-\Phi/K_0},$$
 167

166

are then exactly the same as those from the isokinetic equa- 168 tions of motion given above 169

170 
$$\{\ddot{q} = (F/m) - p \sum (F \cdot \dot{q}/2K_0)e^{+\Phi/2K_0} = (F/m) - \zeta \dot{q}\};$$

$$\zeta = \sum F \cdot \dot{q} / (2K_0).$$

### 172 III. THE HOOVER-LEETE ISOKINETIC THERMOSTAT 173 IN ITS LAGRANGIAN AND HAMILTONIAN 174 FORMS

175 The Hamiltonians corresponding to the isokinetic 176 Gauss'-Principle and Nosé-Hoover velocity-based ap-177 proaches are not the only such means of thermostating equi-178 librium systems. A straightforward application of nonholo-179 nomic Lagrangian mechanics, as well as the familiar 180 Hamiltonian mechanics outlined in Leete's Master's thesis,<sup>6</sup> 181 lead to another type of isokinetic mechanics. These two 182 equivalent forms of mechanics, both of which can be used to 183 keep the *velocity*-based kinetic energy  $K(\dot{q})$  constant, while 184 allowing the momentum-based kinetic energy K(p) to vary, 185 proceed by modifying the  $\{\dot{q}\}$ , rather than the  $\{\dot{p}\}$ , equations. 186 To distinguish the two different kinetic energies we use 187 the notation

188 
$$K(\dot{q}) = \sum \frac{m\dot{q}^2}{2}; \quad K(p) = \sum \frac{p^2}{2m}.$$

189 Let us begin with the Lagrangian case

190 
$$\mathcal{L}_{\text{HL}} \equiv K(\dot{q}) - \Phi + \lambda [K(\dot{q}) - K_0],$$

 where  $K(\dot{q})$  is the *velocity*-based kinetic energy,  $\Sigma(m/2)\dot{q}^2$ , and the Lagrange multiplier  $\lambda$  has the task of maintaining  $K(\dot{q})$  at its initial value,  $K_0$ , as the motion proceeds. Lagrange's equations of motion follow from the usual text- book differentiations of the Lagrangian with respect to the velocities and coordinates:

197 
$$\begin{cases} p = \frac{\partial \mathcal{L}_{\text{HL}}}{\partial \dot{q}} = m\dot{q}(1+\lambda); \\ \dot{p} = m\ddot{q}(1+\lambda) + m\dot{q}\dot{\lambda} = \frac{\partial \mathcal{L}_{\text{HL}}}{\partial q} \equiv F \end{cases}.$$

**199** Now multiply the  $\dot{p}$  equation by  $\dot{q}$  and sum

200 
$$\sum m\ddot{q}(1+\lambda)\dot{q} + \sum m\dot{q}\dot{\lambda}\dot{q} = 0 + 2K_0\dot{\lambda}$$
201 
$$= \sum F \cdot \dot{q} \equiv -\dot{\Phi}.$$

**202** The value of the Lagrange multiplier  $\lambda$  follows:

203 
$$\lambda = \frac{(\Phi_0 - \Phi)}{2K_0} = \frac{\sqrt{4K(p)K_0 - K_0}}{2K_0} = \frac{2K_0(1 + \lambda) - 2K_0}{2K_0} = \lambda,$$

**204** where the last expression, which completes the identity, fol-**205** lows from the Hamiltonian given just below.

206 With Leete's help, Hoover discovered, in 1979, that the 207 *velocity*-based kinetic energy  $K(\dot{q})$  can alternatively be kept 208 constant by using the (constant in time) Hoover-Leete 209 Hamiltonian

$$\mathcal{H}_{\rm HL} \equiv \sum \dot{q}p - \mathcal{L}_{\rm HL} = \sqrt{4K_0K(p) + \Phi - K_0}$$
210

$$=\mathcal{H}_0=\Phi_0+K_0,$$
 211

where  $\Phi$  is again the usual potential energy and  $\Phi_0$  is its 212 initial value. 213

The constancy of the kinetic energy is easy to see. The 214 equations of motion 215

$$\left\{\dot{q} = \frac{\partial \mathcal{H}}{\partial p}; \quad \dot{p} = -\frac{\partial \mathcal{H}}{\partial q}\right\} \longrightarrow \left\{\dot{q} = \frac{p}{m}\sqrt{\frac{K_0}{K(p)}}; \quad \dot{p} = F(q)\right\},$$
216

imply that the velocity-based kinetic energy does not vary 217

$$K(\dot{q}) = \frac{m}{2} \sum \dot{q}^2 = \sum \frac{p^2}{2m} \left[ \frac{K_0}{K(p)} \right] \equiv K_0.$$
 218

At the same time, except in the equilibrium case with a large 219 number of degrees of freedom, there is no guarantee that the 220 momentum-based kinetic energy K(p) is similar in magni- 221 tude to  $K_0$ . But, provided that  $\{m\dot{q}=p\} \rightarrow K(p)=K(\dot{q})=K_0$  ini- 222 tially, Leete's equations of motion do match the Gaussian 223 isokinetic ones to second order 224

$$\left\{ \ddot{q} = \frac{F}{m} - \zeta \dot{q} \right\}; \quad \zeta = \frac{\sum F \cdot \dot{q}}{2K_0},$$
225

through the second derivatives,  $\{\ddot{q}\}$ , but differ in the *third* 226 derivatives, where the mean-squared constraint force in the 227 isokinetic case is less than the Hoover-Leete analog, in ac- 228 cordance with Gauss' Principle of Least Constraint.<sup>1,5</sup> 229

The constraint of constant temperature in Hamiltonian 230 mechanics contradicts the thermodynamic notion that the en- 231 ergy and temperature cannot be varied independently in a 232 system of fixed composition and volume. Both E and T are 233 constants of the motion, using Leete's approach. Provided 234 that the initial conditions are wisely chosen, with E and T 235 corresponding to the *same* thermodynamic state, this ap- 236 proach *can* certainly be used to determine equilibrium prop- 237 erties. We will see, in Sec. VIII, what the (rather strange) 238 consequences of this thermostat are away from equilibrium. 239

### IV. LANDAU-LIFSHITZ' CONFIGURATIONAL 240 THERMOSTAT 241

A (much) more complicated Hamiltonian thermostat, 242 conserving the force-based *configurational* temperature  $T_{con}$ , 243 can be based on straightforward (though quite tedious) holo- 244 nomic Hamiltonian mechanics. A proper configurationally 245 thermostated Hamiltonian conserves not only the Hamil- 246 tonian but also the configurational temperature 247

$$kT_{\rm con} \equiv \sum_{N} F_i^2 / \sum_{N} \nabla_i^2 \mathcal{H}.$$
 248

Here *k* is Boltzmann's constant, which we set equal to unity 249 in the numerical work. A molecular dynamics simulation 250 based on this definition of temperature follows standard 251 Hamiltonian mechanics, as the temperature constraint is just 252 a (complicated) holonomic (coordinates only) constraint. The 253 simplest procedure begins with the system Lagrangian, aug- 254 mented with a Lagrange multiplier  $\lambda$  which constrains 255



FIG. 1. Temperature profile according to Fourier's law, with mean reservoir temperatures of 0.26(-200 < x < -100) and 0.24(0 < x < 100).

**256**  $\{T_{con}, \dot{T}_{con}, \ddot{T}_{con}\}$ . The corresponding Hamiltonian has the **257** form

**258** 
$$\mathcal{H} = \Phi(\{q\}) + K(\{p\}) + \lambda(T_{con} - T_0)$$

 For success, the initial conditions have to be chosen to cor- respond to both the desired temperature,  $T_{\rm con}=T_0$ , and to the condition  $\dot{T}_{\rm con}=0$ . Then, to begin the analytic work, differen- tiate the temperature equation,  $T_{\rm con}(\{q\})=T_0$ , twice with re- spect to time. The two time differentiations, using the chain rule, give first  $\{\dot{q}\}$  and then  $\{\ddot{q}\}$ . By substituting the con-strained equations of motion

**266** 
$$\{m\ddot{q}_i = \dot{p}_i = F_i(\{q\}) - \lambda \nabla_i T_{\rm con}\},\$$

**267** for the  $\{\ddot{q}\}\$ , we obtain the Lagrange multiplier  $\lambda$ . For further **268** details of this calculation, see Sec. IX. The numerical work **269** can be checked by noting that both  $T=T_{\rm con}$  and  $\mathcal{H}$  are con-**270** stants of the motion when the calculation is error free.

# 271 V. AOKI AND KUSNEZOV'S $\varphi^4$ MODEL FOR HEAT 272 CONDUCTION

**273** The  $\phi^4$  model gets its name from the functional form of **274** a quartic "tethering potential,"  $\Phi_{\text{Teth}}$ , which links each par-**275** ticle to a fixed lattice site with a cubic restoring force

$$\Phi_{\text{Teth}} \equiv \sum_{N} \left( \frac{\kappa \delta^4}{4} \right) \rightarrow F_{\text{Teth}} = -\kappa \delta^3.$$

277 In the pedagogical simulations which follow we will choose 278 the tethering force constant equal to unity,  $\kappa = 1$ . Nearest-279 neighbor particle pair interactions in the  $\phi^4$  model are gov-280 erned by Hooke's Law

$$\Phi_{\text{Pair}} = \sum_{\text{NN}} (\mu/2)(|r| - d)^2.$$

276

281

 Aoki and Kusnezov have carried out a variety of pedagogical heat flow simulations for this model in both one and three space dimensions.<sup>17,18</sup> The  $\phi^4$  model has a finite Fourier con- ductivity in one dimension. The numerical work carried out by Aoki and Kusnezov established the temperature depen-dence of the thermal conductivity,  $3/T^{4/3}$ .

**288** For simplicity in the one-dimensional work which we **289** carry out here, we always choose the nearest-neighbor sepa-

ration d equal to unity. We also choose the strength of the 290 Hooke's-Law interaction  $\mu$  and the particle mass equal to 291 unity. As was abundantly demonstrated by Aoki and Kusn- 292 ezov, the combination of a site-based tethering potential with 293 a Hooke's-Law pair potential provides the usual Fourier con- 294 ductivity, with the heat flux proportional to the (sufficiently 295 small) temperature gradient. In what follows we apply four 296 different thermostat constraints to this potential model, using 297 four different dynamical approaches: Nosé-Hoover, Gaussian 298 isokinetic, Hoover-Leete isokinetic, and the Landau-Lifshitz 299 isoconfigurational.

A numerical solution of the heat flow equation for the 301 one-dimensional chain 302

$$\dot{T} = \nabla (3T^{-4/3} \nabla T) \pm \alpha T,$$
303

can be obtained by Runge-Kutta integration, with a rescaling 304 of the temperatures within the reservoir regions at the end of 305 each time step accounting for the source and sink terms  $\pm \alpha T$  306 in the flow equation. Figure 1 shows a fully converged tem- 307 perature profile obtained in this way, using hot and cold tem- 308 peratures of 0.26 and 0.24, with 400 mesh points. The spatial 309 gradient operations were replaced by finite-difference ap- 310 proximations 311

$$(\nabla T)_i = (T_{i+1} - T_{i-1})/2,$$
 312

with appropriate subscript changes accounting for the peri- 313 odic boundary conditions. 314

### VI. RESULTS WITH THE NOSÉ-HOOVER 315 THERMOSTAT 316

With a periodic system composed of four parts, "Hot," "317Newton<sub>1</sub>," "Cold," and "Newton<sub>2</sub>," each part containing *N*/4 318particles, the equations of motion are 319

$$\{m\ddot{q} = F - \zeta_{\text{Hot}}m\dot{q}\}; \quad \dot{\zeta}_{\text{Hot}} = [(K/K_{\text{Hot}}) - 1]/\tau^2,$$
 320

$$\{m\ddot{q} = F - \zeta_{\text{Cold}}m\dot{q}\}; \quad \dot{\zeta}_{\text{Cold}} = [(K/K_{\text{Cold}}) - 1]/\tau^2,$$
 321

322

in the hot and cold regions, and



FIG. 2. Temperature profile for a conventional two-temperature Nosé-Hoover simulation, as described in the text. Here, and in the following figures, the kinetic temperature is indicated by a (quite wiggly) line, while the configurational temperature is indicated by filled circles.



FIG. 3. Temperature profile for a conventional two-temperature Gaussian isokinetic simulation, as described in the text. The kinetic temperature corresponds to the wiggly line; the configurational temperature to the filled circles.

**323** 
$$\{m\ddot{q} = F\},$$

 in the two Newtonian regions. The  $\phi^4$  force  $F_i$  depends upon the three coordinates  $\{x_{i-1}, x_i, x_{i+1}\}$  and also includes the teth- ering force,  $F_{\text{Teth}} = -\kappa (x_i - x_0)^3$ , where  $x_0$  is the lattice site for the *i*th particle in the perfect lattice. The initial velocities and displacements were selected randomly within ranges

**329** 
$$-\Delta_v < \{v\} < +\Delta_v; \quad -\Delta_q < \{\delta q\} < +\Delta_q,$$

 with  $\Delta_v$  or  $\Delta_q$  chosen to reproduce the desired stationary reservoir temperatures, 0.26 and 0.24. The *average* tempera- tures, both kinetic and configurational, from the last half of a 400-particle, 40-million time step simulation,

**334**  $0 < t < 100\ 000\ (equilibration);$ 

**335** 100 000 < t < 200 000( averaging),

336 are shown in Fig. 2, along with a comparison solution of the337 macroscopic heat flow equation. The agreement is quite338 good, as would be expected from Aoki and Kusnezov's339 work.

### **340 VII. RESULTS WITH THE GAUSSIAN ISOKINETIC 341 THERMOSTAT**

342 Simulations with two *Gaussian isokinetic thermostats*, 343 with the same time step, dt=0.005, and the same equilibra-344 tion and averaging intervals, produced similar results, main-345 taining the initial kinetic temperatures throughout. The aver-346 aged profiles, along with a corresponding continuum 347 solution, are shown in Fig. 3. Again the agreement, for both 348 the kinetic and the configurational temperatures, is perfectly 349 satisfactory.

The instantaneous (at time t) values of the total heat added (from time 0 to time t) to the "hot" reservoir as well as that removed from the "cold" reservoir are shown in Fig. 4. The sum of the two necessarily oscillates about zero as tabout zero as tand increases. Straight lines drawn through the data provide an set estimate for the Fourier conductivity in agreement with the affect results of Aoki and Kusnezov's work. The two conventional



FIG. 4. Summed-up values of the heat transferred to the hot and cold Gaussian isokinetic reservoirs for the simulation of Fig. 3.

thermostat types, Nosé-Hoover and Gaussian isokinetic, **357** show that either method can be used to simulate the simple **358** two-temperature heat flow problem. **359** 

## VIII. RESULTS WITH THE HOOVER-LEETE360HAMILTONIAN ISOKINETIC THERMOSTAT361

Results with the Hamiltonian isokinetic thermostats 362

$$\left\{\dot{q}_{\text{Hot/Cold}} = \frac{p}{m} \sqrt{\frac{K_{\text{Hot/Cold}}}{K(p)}}; \quad \dot{p} = F(q)\right\},$$
363

where the fixed hot and cold kinetic temperatures 364  $\langle m\dot{q}^2/k \rangle_{\text{Hot/Cold}}$  are fixed by the initial conditions, but come to 365 differ from the unconstrained momentum-based temperatures 366  $\langle p^2/mk \rangle_{\text{Hot/Cold}}$ . Unlike the Nosé-Hoover and Gaussian isoki- 367 netic profiles, the Hamiltonian-based profiles show *no tem-* 368 *perature gradients at all* (see Fig. 5). Instead there are dis- 369 continuities between the fixed kinetic temperatures and the 370 adjacent constant values of the unconstrained Newtonian re- 371 gions. In the latter regions the kinetic and configurational 372 temperatures equilibrate, and match the unconstrained con- 373 figurational temperature of the thermostated regions. In the 374



FIG. 5. Temperature profiles (kinetic and configurational) for a simulation using constrained Hoover-Leete isokinetic reservoirs. The time step is dt = 0.005, with a total elapsed time of  $10^6$  divided equally between equilibration and averaging portions. The kinetic temperature corresponds to the wiggly line; the configurational temperature to the filled circles.

375 thermostated regions there is no such equilibration.

Evidently there are no heat fluxes in the system either. By constraining not only the energy (through the constant Hamiltonian driving the motion) but also the kinetic temperatures, the system is evidently prevented from supporting heat flow. The Hamiltonian nature of the motion equations all also prevents the formation of the multifractal phase-space all distributions associated with nonequilibrium stationary states.<sup>1,2</sup> A more detailed analysis of the reasons for the failall ure of this Hamiltonian approach could perhaps be based on all analysis of the microscopic equations describing heat flux.<sup>1</sup>

### 387 IX. RESULTS WITH THE LANDAU-LIFSHITZ 388 HAMILTONIAN ISOCONFIGURATIONAL THERMOSTAT

**389** For the special  $\phi^4$  model considered here, with d=1, the **390** force  $F_i$  is

**391** 
$$F_i = -\kappa \delta_i^3 + \mu [\delta_{i+1} - 2\delta_i + \delta_{1-1}].$$

**392** The *pair-potential* contribution to  $\nabla_i^2 \mathcal{H}$  is  $2\mu$ , for the two **393** nearest-neighbor interactions

$$\frac{\partial^2}{\partial x_i^2} (\mu/2) (|x_i - x_{i\pm 1}| - 1)^2 = \mu$$
394

**395** The *tethering* potential contribution is  $3\kappa \delta_i^2 = \nabla_i^2 (\kappa \delta_i^4/4)$ 

$$\sum_{N} \nabla^2 \mathcal{H} = 2N\mu + \sum_{N} 3\kappa \delta^2.$$

**397** The time derivative of the configurational temperature is a **398** quotient of products of single-particle sums

$$\dot{T} = \frac{2\sum F\dot{F}}{\sum (2\mu + 3\kappa\delta^2)} - \frac{6\sum F^2\sum \kappa v\,\delta}{\left[\sum (2\mu + 3\kappa\delta^2)\right]^2},$$

400 where

**401** 
$$\dot{F}_i = \mu [v_{i+1} - 2v_i + v_{i-1}] - 3\kappa \delta_i^2 v_i$$

402 Another time differentiation gives the second derivative

415

- 416
- 417
- 418



 $\lambda = \frac{\ddot{T}_{\text{Newton}}}{\sum \nabla_i T \left( \left[ 2\mu (F_{i+1} - 2F_i + F_{i-1}) - 6\kappa \delta_i^2 F_i \right] - \left[ \sum F^2 \sum 6\kappa F_i \delta_i \right] \right)}$ 

423 A typical profile using two configurational thermostats is 424 shown in Fig. 6. Just as in the Hoover-Leete isokinetic case 425 there is no equilibration of the constrained temperature (here 426 configurational). The unconstrained kinetic temperature 427 equilibrates throughout the system to a value dependent upon



FIG. 6. Temperature profiles (kinetic and configurational) for a simulation using constrained Landau-Lifshitz isoconfigurational reservoirs. The time step is dt=0.002, with a total elapsed time of 500 000 divided equally between equilibration and averaging portions. The kinetic temperature corresponds to the wiggly line; the configurational temperature to the filled circles.

$$\ddot{T}_{\text{Newton}} = \frac{2\sum F\ddot{F} + 2\sum \dot{F}^2}{\sum (2\mu + 3\kappa\delta^2)} + \frac{72\sum F^2 \left[\sum \kappa v \delta\right]^2}{\left[\sum (2\mu + 3\kappa\delta^2)\right]^3} - \frac{24\sum F\dot{F}\sum \kappa v \delta + 6\sum F^2 \left(\kappa \sum \left[a\delta + \sum v^2\right]\right)}{\left[\sum (2\mu + 3\kappa\delta^2)\right]^2}, \quad 404$$

where

$$\ddot{F}_i = \mu [a_{i+1} - 2a_i + a_{i-1}] - \kappa [3\delta_i^2 a_i + 6\delta_i v_i^2].$$
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The *second* derivative,  $\ddot{T}$ , introduces the Lagrange multiplier **407** through the equations of motion **408** 

$$\{a_i = \dot{v}_i = F_i - \lambda \nabla_i T\}.$$

Though complicated, the resulting equation for  $\lambda$  is linear, 410 and gives an explicit expression for the Lagrange multiplier 411 that stabilizes T 412

the initial conditions. Evidently the Landau-Lifshitz con- 428 strained configurational thermostat is quite useless for non- 429 equilibrium work. 430

### X. CONCLUSIONS

The results obtained with the Gaussian isokinetic and 432 Nose-Hoover thermostats are quite consistent with the past 433 history of their use over the last quarter century.<sup>1,2,6,7</sup> For 434 these familiar thermostats there was no problem in reaching 435

436 a nonequilibrium steady state with temperature profiles and 437 heat fluxes very similar to those expected from macroscopic **438** hydrodynamics using Fourier's law. Although the simple  $\phi^4$ 439 model may strike the reader as rather special we believe that 440 the results obtained using it are typical of realistic applica-441 tions. The lack of momentum conservation in this model, due 442 to the tethering potential, though it certainly eliminates the 443 low-frequency divergence of the one-dimensional heat con-444 ductivity, actually makes the one- and two-dimensional mod-445 els more closely resemble their more physical three-446 dimensional cousins. We expect that the configurational 447 analog<sup>14</sup> of the Nosé-Hoover kinetic thermostat, discussed in 448 Sec. II, would provide very similar results. It would be par-449 ticularly interesting to explore the differences between the 450 tensor forms of the kinetic and configurational temperatures, 451 by applying these ideas to strong shockwaves.

On the other hand, the results obtained with purely 452 453 Hamiltonian mechanics, using either the Hoover-Leete con-454 strained isokinetic or the Landau-Lifshitz constrained con-455 figurational thermostats were both unexpected and thought-456 provoking. The main message of the current work is that 457 unadulterated Hamiltonian mechanics is simply unsuitable to 458 steady-state nonequilibrium simulations. The longtime re-459 sults, using Hamiltonian mechanics, are tied to both the en-460 ergy and the phase volume of the initial conditions, in a way 461 simply inconsistent with the known phase-space contraction 462 typical of nonequilibrium systems. From the æsthetic stand-463 point, we believe that the canonical-ensemble derivations of 464 Landau and Lifshitz<sup>12</sup> and Travis and Braga<sup>14</sup> are not only 465 clearer, but also more useful and stimulating than is Rugh's **466** alternative microcanonical approach.<sup>13</sup>

467 Both the Hamiltonian approaches stably constrained 468 their target temperatures in the two reservoir regions. Never-469 theless temperature differences led to no net heat flow and 470 the computed temperature profiles bore no resemblance to 471 the predictions of hydrodynamics. The kinetic and configu-472 rational temperatures, though *unequal* in the constrained res-473 ervoirs

 $T_{\rm kin} \neq T_{\rm con}$  [reservoirs], 474

475 were equal (to yet a third nearly constant temperature) in the 476 unconstrained bulk Newtonian regions

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 $T_{\rm kin} = T_{\rm con}$  [bulk]. 477

These results emphasize the fact that nonHamiltonian dy- 478 namics is absolutely necessary to the realistic modeling of 479 nonequilibrium systems. A direct proof/understanding of the 480 failure of the Hamiltonian systems to show a nonvanishing 481 heat flux would be most welcome. 482

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